

Observed: EVPA \perp jet axis in most BL Lac jets
but tend to be more often oblique in quasar jets

B. Faraday rotation e^-p^+ plasmas only

$$\Delta\chi = \frac{e^3}{2\pi m^2 c^2} \nu^{-2} \int_0^S n_e \vec{B} \cdot d\vec{S} = 7.3 \times 10^4 \nu_{\text{GHz}}^{-2} \int_0^{S_{\text{pc}}} n_e B_{\text{G.O.S.}} ds \text{ radians}$$

$$n_e = n_e(\text{thermal}) + \frac{n_e(\text{relativistic})}{\langle \gamma \rangle^2}$$

If $n_e B_{\text{G.O.S.}}$ varies across source, + $\Delta\chi$ is large, effect causes depolarization from differential $\Delta\chi$'s

C. Circular Polarization

Tends to be at level $\Pi_{\text{circ}} \approx \gamma^{-1}(\nu) \frac{B_{\text{unif}}^2}{B_{\text{unif}}^2 + B_{\text{rand}}^2}$ for optically thin source
typically $\Pi_{\text{circ}} \lesssim 0.1\%$ expected
Legg + Westfold (1968, ApJ, 154, 499)
 $\gamma(\nu) = 19 \nu_{\text{GHz}}^{\frac{1}{2}} B^{-\frac{1}{2}} \left(\frac{1+z}{5}\right)^{\frac{1}{2}}$

Observed (Wardle, Homan, et al.): up to $\sim 1\%$ (e.g., 3C 279)
→ Conversion of linear to circular polarization in optically thick region (plasma effect)

D. SSC Polarization (Celotti + Matt 1994, MNRAS, 268, 451)
linear $\Pi \sim 50\%$ of synchrotron Π , same χ

V. Matter Content of Jets: e^-p^+ or e^-e^+ ? or both \sim equal?

Galactic cosmic-rays: e^-p^+ , $U(p^+) \sim 100 U(e^-)$

e^-p^+ advantages: more momentum, inertia → harder to disrupt, $\Gamma = \text{const.}$
after initial acceleration, as observed in many objects ($\beta_{\text{app}} \approx \text{const.}$)

e^-e^+ advantages: high-E processes make e^-e^+ pairs, not much rest mass
so easier to accelerate

Attempts to decide issue are all flawed or not definitive

- Power needed to feed lobes - depends on value of δ_{min} of e^-e^+
- " " " create high-E emission - uses δ -function approx. for SSC that is incorrect for X-ray emission
- Circular polarization: Wardle + Homan claim low $\Delta\chi$ needed while high optical depth \Rightarrow is present → mostly (80%) e^-e^+ plasma
→ not true according to calculations by Ruszkowski + Begelman (2002, ApJ, 573, 485)